

**Table 1.** Median values and 68% confidence interval for OGLE-TR-1072.

Parameter	Units	Values
Stellar Parameters:		
$M_*$	Mass ( $M_\odot$ )	$0.97^{+0.16}_{-0.15}$
$R_*$	Radius ( $R_\odot$ )	$1.137^{+0.069}_{-0.068}$
$R_{*,SED}$	Radius <sup>1</sup> ( $R_\odot$ )	$1.137^{+0.078}_{-0.073}$
$L_*$	Luminosity ( $L_\odot$ )	$1.74^{+0.62}_{-0.41}$
$F_{Bol}$	Bolometric Flux (cgs)	$0.000000000249^{+0.0000000000078}_{-0.0000000000052}$
$\rho_*$	Density (cgs)	$0.94^{+0.22}_{-0.18}$
$\log g$	Surface gravity (cgs)	$4.317^{+0.073}_{-0.077}$
$T_{eff}$	Effective Temperature (K)	$6190^{+570}_{-380}$
$T_{eff,SED}$	Effective Temperature <sup>1</sup> (K)	$6200^{+550}_{-400}$
[Fe/H]	Metallicity (dex)	$-0.29^{+0.39}_{-2.3}$
[Fe/H] <sub>0</sub>	Initial Metallicity <sup>2</sup>	$-0.20^{+0.32}_{-1.8}$
Age	Age (Gyr)	$6.4^{+4.9}_{-4.5}$
EEP	Equal Evolutionary Phase <sup>3</sup>	$403^{+24}_{-63}$
$A_V$	V-band extinction (mag)	$1.08^{+0.31}_{-0.26}$
$\sigma_{SED}$	SED photometry error scaling	$10.4^{+1.6}_{-1.3}$
$\varpi$	Parallax (mas)	$0.671^{+0.042}_{-0.041}$
$d$	Distance (pc)	$1490^{+97}_{-87}$
Planetary Parameters:		
		b
$P$	Period (days)	$1.6133902 \pm 0.0000014$
$R_P$	Radius ( $R_J$ )	$0.993^{+0.070}_{-0.068}$
$M_P$	Mass <sup>4</sup> ( $M_J$ )	$49^{+24}_{-29}$
$T_C$	Time of conjunction <sup>5</sup> (BJD <sub>TDB</sub> )	$2455377.4959 \pm 0.0015$
$T_T$	Time of minimum projected separation <sup>6</sup> (BJD <sub>TDB</sub> )	$2455377.4959 \pm 0.0015$
$T_0$	Optimal conjunction Time <sup>7</sup> (BJD <sub>TDB</sub> )	$2456919.89690 \pm 0.00069$
$a$	Semi-major axis (AU)	$0.0271 \pm 0.0014$
$i$	Inclination (Degrees)	$84.5^{+2.0}_{-1.6}$
$T_{eq}$	Equilibrium temperature <sup>8</sup> (K)	$1930^{+190}_{-130}$
$\tau_{circ}$	Tidal circularization timescale (Gyr)	$1.32^{+1.1}_{-0.87}$
$K$	RV semi-amplitude <sup>4</sup> (m/s)	$8400^{+4000}_{-4800}$
$R_P/R_*$	Radius of planet in stellar radii	$0.0898 \pm 0.0019$
$a/R_*$	Semi-major axis in stellar radii	$5.13^{+0.37}_{-0.35}$
$\delta$	$(R_P/R_*)^2$	$0.00807^{+0.00035}_{-0.00034}$
$\delta_I$	Transit depth in I (fraction)	$0.00883 \pm 0.00033$
$\delta_V$	Transit depth in V (fraction)	$0.00939^{+0.00049}_{-0.00041}$
$\tau$	Ingress/egress transit duration (days)	$0.0105^{+0.0019}_{-0.0016}$
$T_{14}$	Total transit duration (days)	$0.0986^{+0.0021}_{-0.0019}$

Table 1 continued on next page

Table 1 (continued)

Parameter	Units	Values	
$T_{FWHM}$ ..	FWHM transit duration (days) .....	0.0880 ± 0.0017	
$b$ .....	Transit Impact parameter .....	0.49 <sup>+0.10</sup> <sub>-0.15</sub>	
$\delta_{S,2.5\mu m}$ ..	Blackbody eclipse depth at 2.5 $\mu m$ (ppm) .....	670 <sup>+130</sup> <sub>-100</sub>	
$\delta_{S,5.0\mu m}$ ..	Blackbody eclipse depth at 5.0 $\mu m$ (ppm) .....	1390 <sup>+160</sup> <sub>-140</sub>	
$\delta_{S,7.5\mu m}$ ..	Blackbody eclipse depth at 7.5 $\mu m$ (ppm) .....	1720 <sup>+160</sup> <sub>-150</sub>	
$\rho_P$ .....	Density <sup>4</sup> (cgs) .....	62 <sup>+37</sup> <sub>-38</sub>	
$\log g_P$ .....	Surface gravity <sup>4</sup> .....	5.09 <sup>+0.19</sup> <sub>-0.40</sub>	
$\Theta$ .....	Safronov Number .....	2.8 <sup>+1.5</sup> <sub>-1.7</sub>	
$\langle F \rangle$ .....	Incident Flux (10 <sup>9</sup> erg s <sup>-1</sup> cm <sup>-2</sup> ) .....	3.17 <sup>+1.4</sup> <sub>-0.77</sub>	
$T_P$ .....	Time of Periastron (BJD <sub>TDB</sub> ) .....	2455377.4959 ± 0.0015	
$T_S$ .....	Time of eclipse (BJD <sub>TDB</sub> ) .....	2455376.6892 ± 0.0015	
$T_A$ .....	Time of Ascending Node (BJD <sub>TDB</sub> ) .....	2455378.7059 ± 0.0015	
$T_D$ .....	Time of Descending Node (BJD <sub>TDB</sub> ) .....	2455377.8992 ± 0.0015	
$V_c/V_e$ .....	.....	1.00	
$M_P \sin i$ ..	Minimum mass <sup>4</sup> ( $M_J$ ) .....	49 <sup>+24</sup> <sub>-29</sub>	
$M_P/M_*$ .....	Mass ratio <sup>4</sup> .....	0.048 <sup>+0.025</sup> <sub>-0.028</sub>	
$d/R_*$ .....	Separation at mid transit .....	5.13 <sup>+0.37</sup> <sub>-0.35</sub>	
$P_T$ .....	A priori non-grazing transit prob .....	0.177 <sup>+0.013</sup> <sub>-0.012</sub>	
$P_{T,G}$ .....	A priori transit prob .....	0.212 <sup>+0.016</sup> <sub>-0.015</sub>	
Wavelength Parameters:		I	V
$u_1$ .....	linear limb-darkening coeff .....	0.234 <sup>+0.067</sup> <sub>-0.062</sub>	0.379 <sup>+0.087</sup> <sub>-0.071</sub>
$u_2$ .....	quadratic limb-darkening coeff .....	0.290 ± 0.052	0.294 <sup>+0.054</sup> <sub>-0.058</sub>
Transit Parameters:		OGLE UT 2010-06-29 (I)	OGLE UT 2010-06-29 (V)
$\sigma^2$ .....	Added Variance .....	0.00002844 <sup>+0.00000049</sup> <sub>-0.00000048</sub>	0.0000358 <sup>+0.0000057</sup> <sub>-0.0000047</sub>
$F_0$ .....	Baseline flux .....	1.000272 <sup>+0.000058</sup> <sub>-0.000059</sub>	1.00017 ± 0.00054

See Table 3 in Eastman, J. et al., 2019, arXiv:1907.09480 for a detailed description of all parameters

<sup>1</sup>This value ignores the systematic error and is for reference only

<sup>2</sup>The metallicity of the star at birth

<sup>3</sup>Corresponds to static points in a star's evolutionary history. See §2 in Dotter, A., 2016, ApJS, 222, 8

<sup>4</sup>Uses measured radius and estimated mass from Chen, J., & Kipping, D. 2017, ApJ, 834, 17

<sup>5</sup>Time of conjunction is commonly reported as the "transit time"

<sup>6</sup>Time of minimum projected separation is a more correct "transit time"

<sup>7</sup>Optimal time of conjunction minimizes the covariance between  $T_C$  and Period

<sup>8</sup>Assumes no albedo and perfect redistribution